

NOTES: Radiation properties

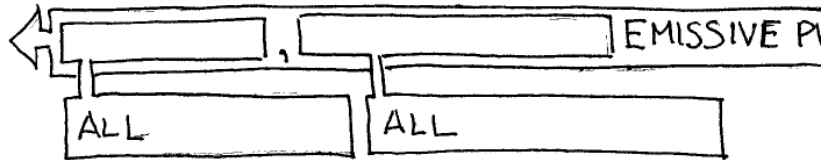
RADIATION PROPERTIES

FOR BLACKBODIES



$$E_b(T) = \sigma T^4$$

$$\equiv \frac{[\quad]}{[\quad]}$$

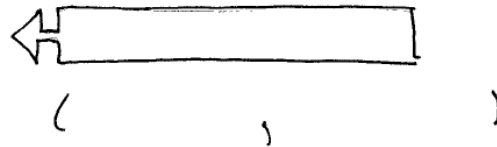


A REAL BODY EMITS LESS

$$E(T) = \epsilon E_b = \epsilon \sigma T^4$$

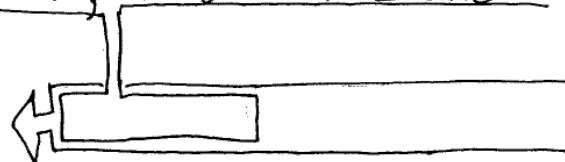
-OR-

$$\epsilon \equiv \frac{E(T)}{E_b(T)}$$



@ A PARTICULAR WAVELENGTH, PER UNIT WAVELENGTH

$$\epsilon_\lambda \equiv \frac{E_\lambda(T)}{E_{b,\lambda}(T)}$$



IF A SURFACE IS

_____ , ITS PROPERTIES ARE INDEPENDENT of

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CAN ALSO DEFINE

ϵ_θ :

$\epsilon_{\lambda,\theta}$:

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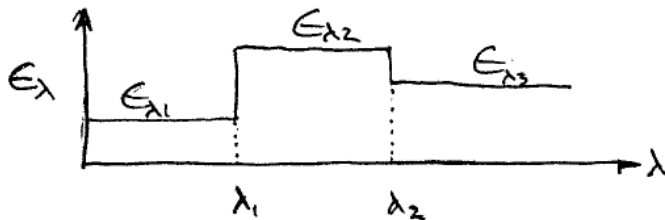
FOR A NON-GRAY SURFACE

$$E = E_{b,\lambda} = \sigma T^4$$

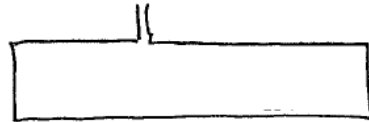
SO

$$E(T) = \frac{\int_0^{\infty} \epsilon_{\lambda} E_{b,\lambda} d\lambda}{\sigma T^4}$$

FOR ϵ_{λ} THAT VARIES IN A STEP-LIKE FASHION; i.e.



$$E = \frac{\int_0^{\lambda_1} \epsilon_{\lambda_1} E_{b,\lambda} d\lambda}{\sigma T^4} + \epsilon_{\lambda_2} \frac{\int_{\lambda_1}^{\lambda_2} E_{b,\lambda} d\lambda}{\sigma T^4} + \dots$$



THUS

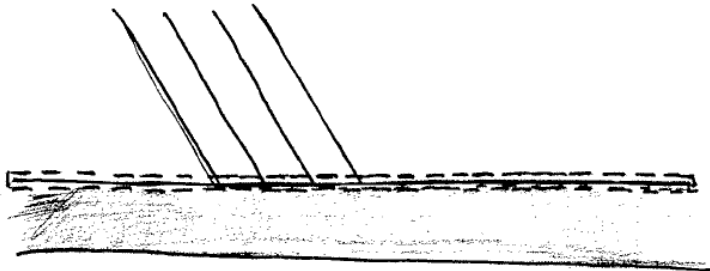
$$E = \epsilon_{\lambda_1} + \epsilon_{\lambda_2} + \dots$$

! LIST THE ASSUMPTIONS!

NOTES: Radiation properties

OTHER PROPERTIES

INCIDENT RADIATION ()



$$\alpha = \frac{\quad}{\text{INCIDENT}} = \quad \leftarrow \text{[]} *$$

$$\rho = \frac{\quad}{\text{INCIDENT}} = \quad \leftarrow \text{[]} *$$

$$\tau = \frac{\quad}{\text{INCIDENT}} = \quad \leftarrow \text{[]} *$$

CONS. of ENERGY ON THIS SURFACE REQUIRES

$$G + G + G = G$$

-OR-



IF A SURFACE IS _____

$$\tau = 0 \Rightarrow$$

* NOTE THAT THESE PROPS NOT ONLY DEPEND ON THE SURFACE, BUT ALSO THE SOURCE & THE IRRADIATION!!!

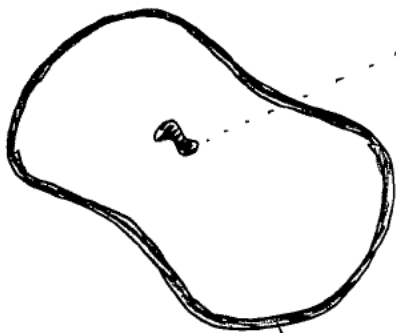
NOTES: Radiation properties

IF SOURCE of G IS A BLACKBODY & α VARIES STEPWISE

$$\alpha = \underbrace{\alpha_{\lambda_1} \int_0^{\lambda_1} E_{\lambda} d\lambda}_{\sigma T_{SOURCE}^4} + \alpha_{\lambda_2} \int_{\lambda_1}^{\lambda_2} E_{\lambda} d\lambda + \dots$$

$$= \alpha_{\lambda_1} + \dots$$

KIRCHOF'S LAW



CONS. of ENERGY

$$\frac{dE_{sys}}{dt} =$$

LARGE, ISOTHERMAL ENCLOSURE @

(\therefore)



WARNING!

IF $T_{SOURCE} \gg T$
OR
 $T_{SOURCE} \ll T$

$\neq !!$

* FOR RESTRICTIONS ON $\alpha_{\lambda, \theta} = \epsilon_{\lambda, \theta}$